



WELCOME TO T.E.I. CRETE Department of Electronic Engineering Faculty of Applied Sciences Centre for Plasma Physics & Lasers (CPPL) http://www.cppl.teicrete.gr/









"In principle our knowledge is limited as it is in the far interior of matter" , 4th century bc





PLASMA & LASER TECHNOLOGY ENJOY THE JOURNAY







Introduction to plasma physics



Definition of Plasma, the "fourth state of matter"

Particle motions in a plasma

Plasma as a fluid

Waves in plasmas

Plasma equilibrium

Kinetic theory











- 1. Introduction to Plasma Physics and Controlled Fusion, Volume 1: Plasma Physics, F.F. Chen NY, Plenum Press, 1984
- 2. Plasma Physics: An Introductory Course, edited by R.Dendy, Cambridge, Cambridge University Press, 1993
- *3. The physics of laser-plasma interactions, W.L.Kruer, Addison-Wesley, 1988*
- 4. Basic Space Plasma Physics, Baumjohann and Treumann, Imperial College Press, London 1997







Plasma is a partially or completely ionised medium, presents a collective behavior (unlike the ideal gas). It is considered as the fourth state of matter (unlike the ideal gas)

It is named "plasma" «πλάσμα» by I. Langmuir and L. Tonks Plasma should not contain more than 1% neutral atoms or molecules











• It is a magnificent material

It is very selfish, does not like external influences but likes group work

It has mind on its own









Plasma is the "fourth state" of matter

Consists of electrons, ions and neutral atoms

Index of refraction < 1

99% of matter of the visible Universe is in plasma state









is any ionised gas plasma?









Definition of plasma

A quasineutral gas of charged and neutral particles which exhibits collective behavior Quasineutral:

The plasma is neutral enough so that one can take n_i~n_e~n

Collective:

Plasma charges move around and generate local concentrations of positive and negative charge and therefore electric fields. Also, motion of charges generates currents, thus magnetic fields. These fields affect the motion of other particles far away.







Debye shielding







Debye shielding

$$\frac{d^2\phi}{dx^2} = \frac{e^2 n_{\infty}}{\varepsilon_o k T_e} \phi, \lambda_D \equiv \left(\frac{\varepsilon_o k T_e}{n e^2}\right)^{1/2}$$



 λ_D Debye length



The plasma parameter

Number of particles in the Debye sphere

$$N_D = n \frac{4}{3} \pi \lambda_D^3 = 1.38 \times 10^6 \frac{T^{3/2}}{n^{1/2}}, T({}^o K)$$

Collective behavior requires:





Plasma	n_e	T	B	λ_D	N_D	ω_p	$ u_{ee}$	ω_c	r_L
	(m^{-3})	(K)	(T)	(m)		(s^{-1})	(s^{-1})	(s^{-1})	(m)
Gas discharge	10^{16}	10^{4}		10^{-4}	10^{4}	10^{10}	10^{5}		
Tokamak	10^{20}	10^{8}	10	10^{-4}	10^{8}	10^{12}	10^{4}	10^{12}	10^{-5}
Ionosphere	10^{12}	10^{3}	10^{-5}	10^{-3}	10^{5}	10^{8}	10^{3}	10^{6}	10^{-1}
Magnetosphere	10^{7}	10^{7}	10^{-8}	10^{2}	10^{10}	10^{5}	10^{-8}	10^{3}	10^{4}
Solar core	10^{32}	10^{7}		10^{-11}	1	10^{18}	10^{16}		
Solar wind	10^{6}	10^{5}	10^{-9}	10	10^{11}	10^{5}	10^{-6}	10^{2}	10^{4}
Interstellar medium	10^{5}	10^{4}	10^{-10}	10	10^{10}	10^{4}	10^{-5}	10	10^{4}
Intergalactic medium	1	10^{6}		10^{5}	10^{15}	10^{2}	10^{-13}		

Source: Chapter 19: *The Particle Kinetics of Plasma* http://www.pma.caltech.edu/Courses/ph136/yr2004/



Plasma Criteria

 $\lambda_D << L$

 $N_{D} >> 1$

$\omega \tau > 1$

- ω is the frequency of typical plasma oscillations
- τ is the mean time between collisions with neutral atoms

$$\begin{array}{ll} n_{o} = \ 10^{20} \ m^{-3} \,, \ T = \ 1 \ keV, & \lambda_{D} = \ 20 \ \mu m \\ n_{o} = \ 10^{8} m^{-3} \,, & T = \ 10 \ eV, & \lambda_{D} = \ 3 \ m \end{array}$$



For a plasma- Saha equation



i.e. hydrogen plasma in thermodynamic equilibrium

$$N_e = N_p \quad N_n$$

$$n_e = n_p n_n$$

Total density:

$$n = n_n + n_e = n_n + n_p$$

gas temperature: $T^{o}K$ ionisation energy: E_{i}



Megh Nad Saha "Saha equation"

$$\frac{n_i}{n_n} = \frac{\left(2\pi m_e kT\right)^{3/2}}{nh^3} e^{-\frac{E_i}{kT}} = 2.4x10^{15} \frac{T^{3/2}}{n_i} e^{-\frac{E_i}{kT}}$$

T=300°K, E_i=14,5eV (nitrogen), n_n=3x10²⁵m⁻³

$$\frac{n_i}{n_n} \approx 10^{-122}$$





	Kind of plasma	Density n [cm-3	B] Temperature T [K]
and a first	Electrical discharges	~10 ¹⁰	~10,000
	Earth's ionosphere	~10 ⁶	~1,000
	Solar wind	~10	Te ~ 500,000, Ti ~ 100,000
	Solar corona	~10 ⁹	~10,000,000
	Tokamak	~10 ¹⁵	~100,000,000
	White dwarfs	~10 ³⁰	~10,000
	Hydrogen bomb	~10 ³⁰	~100,000,000,000
	Interplanetary space	~10 ⁻² -1	~100
	Laser plasma		







Plasma frequency

Η εξίσωση κίνησης των ηλεκτρονίων είναι:

$$\sigma = n_o e \delta \chi$$
$$m_e \frac{d^2 \delta \chi}{dt^2} = -eE = -\frac{n_o e^2 \delta \chi}{\varepsilon_o} \qquad E = \frac{\sigma}{\varepsilon_o} = \frac{V}{d}$$

Τα ηλεκτρόνια εκτελούν <u>αρμονική ταλάντωση</u> γύρω από την θέση ισορροπίας με συχνότητα :

$$\omega_p = \left(\frac{n_0 e^2}{m\varepsilon_0}\right)^{1/2}$$

Αν θεωρήσουμε και την κίνηση των ιόντων τότε

$$\omega_p = \left(\frac{n_0 e^2}{m_e \varepsilon_0} + \frac{n_0 e^2}{m_i \varepsilon_0}\right)^{1/2}$$

η συχνότητα πλάσματος είναι :



Some important plasma parameters

Plasma frequency
$$\omega_{pe} = \sqrt{\frac{n_e e^2}{\varepsilon_o m_e}} = 5.64 \times 10^4 n_e^{1/2} rad \, \mathrm{sec}^{-1}$$

Critical density
$$n_c = \frac{\omega^2 m_e \varepsilon_o}{e^2} = 1.113 \times 10^{21} \left(\frac{1\mu m}{\lambda}\right)$$

Debye length

$$\lambda_{De} = \left(\frac{\varepsilon_o k_B T_e}{n_e e^2}\right)^{1/2} = 2.35 \times 10^{-8} \left(\frac{T_e}{1eV}\right)^{1/2} \left(\frac{10^{21}}{n_e (cm^{-3})}\right)^{1/2} cm$$

2

 cm^{-3}

Debye number
$$N_D = n \frac{4}{3} \pi \lambda_D^3 = 1.38 \times 10^6 T^{3/2} / n^{1/2}$$

Average number of electrons in a plasma contained within a Debye sphere

In a real plasma the positions and motions of the particles (electrons, ions and neutral atoms) are determined by randomness due to thermal effects and unexpected internal or external perturbations. The E and B fields are therefore also determined by the motions of the particles in the plasma and the currents that are generated due to the external or internal perturbations.

Since the density of the plasma can vary from $\sim 10^{12}$ particles/cm³ up to $\sim 10^{24}$ particles/cm³, it is obvious that in order to study the plasma dynamics it is really **almost impossible** to follow each particle's trajectory in time and its interaction with the other particles and the time varying generated **E** and **B** fields.

Fortunately, most of the plasma phenomena for experimentally generated plasmas (i.e. using lasers or pulsed power devices such as Z or X-pinches or Tokamak, can be well described using fluid mechanics physics such as the plasma is fluid*. In this case the behaviour and identity of the individual particles is not taken into account and instead, the motion of fluid elements using the fluid equations adapted to the plasma conditions (and the fact that the fluid contains charged particles as well as **E** and **B** fields) are implemented.



a few orders of magnitude!!



The Plasma as a Fluid - Maxwell's equations

In equations (1) and (4) ρ_f and J_f are the "free" charge and current density. The "bound charge" and current densities arising from the polarisation and magnetisation of the plasma (like in a dielectric medium) are included in the quantities **D** and **H**, which are the <u>electric displacement field</u> and magnetic field **H** (same name as the magnetic field **B**) respectively.

The total charge density is $\rho_q = \rho_f + \rho_b$. So in a plasma the Gauss law would write:

$$\nabla \cdot \boldsymbol{D} = \rho_{f} \Leftrightarrow \nabla \cdot (\varepsilon_{o}\boldsymbol{E} + \boldsymbol{P}) = \varepsilon_{o}(1 + \chi_{\varepsilon})\boldsymbol{E}$$
$$\boldsymbol{D} = \varepsilon \boldsymbol{E} = \varepsilon_{o}\boldsymbol{E} + \boldsymbol{P}$$
$$\boldsymbol{P} = \varepsilon_{o}\chi_{\varepsilon}\boldsymbol{E} \qquad \varepsilon = (1 + \chi_{\varepsilon})\varepsilon_{o}$$
$$\boldsymbol{H} \equiv \frac{\boldsymbol{B}}{\mu_{0}} - \boldsymbol{M}.$$

$$\boldsymbol{B} = \boldsymbol{\mu}_o(\boldsymbol{H} + \boldsymbol{M})$$

Convective derivative

Definition of Convective Derivative:

A derivative taken with respect to a moving coordinate system, also called the Langrangian derivative, substantive derivative, or Stokes derivative. It is given by:

$$\frac{D}{Dt} = \frac{\partial}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla}$$

where is the gradient operator and v is the velocity of the fluid. This type of derivative is especially useful in the study of fluid mechanics. When applied to v,

$$\frac{D\boldsymbol{v}}{Dt} = \frac{\partial \boldsymbol{v}}{\partial t} + (\boldsymbol{\nabla} \times \boldsymbol{v}) \times \boldsymbol{v} + \boldsymbol{\nabla}(\frac{1}{2}\boldsymbol{v}^2)$$

$$(\boldsymbol{A}\cdot\boldsymbol{\nabla})\boldsymbol{A}=\boldsymbol{\nabla}(\frac{1}{2}\boldsymbol{A}^2)-\boldsymbol{A}\times(\boldsymbol{\nabla}\times\boldsymbol{A})$$

If there are no collisions and no thermal motion, all particles in a fluid move together. The average velocity of the fluid in an element equals the individual particle velocity.

Conservation of matter

The total number of particles can be altered only if there is a net flux of particles across the surface S which bounds the volume V

Particle flux density: J = nv

Divergence theorem:

hly if there is
ch bounds the
$$\int_{V} (\nabla \cdot A) dV = \oint_{S} A \cdot dS$$

$$\frac{\partial N}{\partial t} = \int_{V} \frac{\partial n}{\partial t} dV = -\oint_{S} n\boldsymbol{v} \cdot d\boldsymbol{S} = -\int_{V} \nabla \cdot (n\boldsymbol{v}) dV \text{ for any volume V} \implies$$

$$\frac{\partial n}{\partial t} + \boldsymbol{\nabla} \cdot (n\boldsymbol{\upsilon}) = 0$$

Equation of continuity for each specie of the plasma

Equation of state

 $p = Cn^{\gamma}$ γ is the ratio of the specific heats C_p/C_v $\gamma = \frac{2+N}{N}$

N is the number of degrees of freedom



prove this

isothermal change $\gamma=1$ adiabatic/isotropic 3 degrees of freedom $\gamma=5/3$ adiabatic 1 degree of freedom $\gamma=3$ adiabatic 2 degree of freedom $\gamma=2$

isothermal compression:

$$\boldsymbol{\nabla} p = \boldsymbol{\nabla}(nKT) = KT\boldsymbol{\nabla} n$$

ideal gas:
$$p = nKT$$

 $n = \frac{N}{V}$ particle density

Convective derivative

In the absence of collisions and thermal motion, the fluid equation could be obtained by multiplying the equation of motion by the density of the species n

$$mn\frac{\partial \boldsymbol{\upsilon}}{\partial t} = qn\left(\boldsymbol{E} + \boldsymbol{\upsilon} \times \boldsymbol{B}\right)$$

In eq. of motion, time derivative is taken at the position of particles (remember the non uniform E - finite Larmor-radius treatment). But we are keen to have an equation for fluid elements fixed in space, thus need a transformation to variables in a fixed frame. Assume G(x,t) to be any variable of fluid.

$$\frac{dG(x,t)}{dt} = \frac{\partial G}{\partial t} + \frac{\partial G}{\partial x}\frac{dx}{dt} = \frac{\partial G}{\partial t} + \upsilon_x \frac{\partial G}{\partial x} \equiv \frac{DG}{Dt}$$
Change at a fixed point
Change at a fixed point
$$Change of G as the observer moves
with the fluid into a region in which G
is different
$$\upsilon \cdot \nabla \text{ is a scalar differential operator}$$$$

If there are no collisions and no thermal motion, all particles in a fluid move together. The average velocity of the fluid in an element equals the individual particle velocity.

For a plasma G=v the fluid velocity:

$$mn(\left(\frac{\partial \boldsymbol{v}}{\partial t} + (\boldsymbol{v} \cdot \boldsymbol{\nabla})\boldsymbol{v}\right) = qn\left(\boldsymbol{E} + \boldsymbol{v} \times \boldsymbol{B}\right) \qquad \frac{\partial \boldsymbol{v}}{\partial t} = \begin{array}{c} \text{Time derivative in a} \\ \text{fixed frame} \end{array}$$

Thermal motions:

If thermal motions are considered, a pressure force has to be included in the equation due to the random motion of the particles in the plasma fluid.



Lets consider for simplicity only the x component of the motion. The Plasma as a Fluid Ordinary fluid dynamics Navier – Stokes equation

(1)
$$\rho(\frac{\partial \boldsymbol{v}}{\partial t} + (\boldsymbol{v} \cdot \boldsymbol{\nabla})\boldsymbol{v} = -\boldsymbol{\nabla}p + \rho \boldsymbol{v} \boldsymbol{\nabla}^2 \boldsymbol{v} \longleftarrow$$

Viscosity term, Viscosity term, v is the kinematic viscosity coefficient

Comparison with the plasma motion equation:

(2)
$$mn(\frac{\partial \boldsymbol{v}}{\partial t} + (\boldsymbol{v} \cdot \boldsymbol{\nabla})\boldsymbol{v} = qn(\boldsymbol{E} + \boldsymbol{v} \times \boldsymbol{B}) - \boldsymbol{\nabla} \cdot \mathbf{P} - \frac{mn(\boldsymbol{v} - \boldsymbol{v}_o)}{\tau}$$

 $\rho v \nabla^2 v \rightarrow$ Viscosity term corresponds to the collisional part in the fluid eq.

Equation (1) describes a collisional fluid with frequent collisions between particles.

Equation 2 was derived without "collision rate" definition **between plasma species**, but eq. (2) indeed can describe the plasma species. Since we used the Maxwellian velocity distribution <u>we implicitly considered collisions</u>.

The complete set of equations

$$\nabla \cdot \boldsymbol{E} = \frac{\rho_q}{\varepsilon_o} = \frac{n_i q_i + n_e q_e}{\varepsilon_o}$$
$$\nabla \times \boldsymbol{E} = -\frac{\partial \boldsymbol{B}}{\partial t}$$
$$\nabla \times \boldsymbol{B} = \mu_o (\boldsymbol{j} + \varepsilon_o \frac{\partial \boldsymbol{E}}{\partial t}) =$$
$$= \mu_o (n_i q_i \boldsymbol{v}_i + n_e q_e \boldsymbol{v}_e + \varepsilon_o \frac{\partial \boldsymbol{E}}{\partial t})$$
$$\nabla \cdot \boldsymbol{B} = 0$$

$$p_{j} = C_{j} n_{j}^{\gamma} j = i, e$$

$$\frac{\partial n_{j}}{\partial t} + \nabla \cdot (n_{j} \boldsymbol{v}_{j}) = 0$$

$$\boldsymbol{j} = n_i q_i \boldsymbol{v}_i + n_e q_e \boldsymbol{v}_e$$

$$\rho_q = n_i q_i + n_e q_e$$

All charges are included in $\rho_q,$ bound & free

$$m_{j}n_{j}\left(\frac{\partial \boldsymbol{v}_{j}}{\partial t} + (\boldsymbol{v}_{j} \cdot \boldsymbol{\nabla})\boldsymbol{v}_{j}\right) = q_{j}n_{j}\left(\boldsymbol{E} + \boldsymbol{v}_{j} \times \boldsymbol{B}\right) - \boldsymbol{\nabla}\boldsymbol{p}_{j} \quad j=i,e$$





Waves in plasmas





<u>Waves in plasmas</u>

$$\nabla \cdot \boldsymbol{E} = \frac{\rho}{\varepsilon_o} = \frac{n_i q_i + n_e q_e}{\varepsilon_o}$$
$$\nabla \times \boldsymbol{E} = -\frac{\partial \boldsymbol{B}}{\partial t}$$
$$\nabla \times \boldsymbol{B} = \mu_o (\boldsymbol{j} + \varepsilon_o \frac{\partial \boldsymbol{E}}{\partial t}) =$$
$$= \mu_o (n_i q_i \boldsymbol{v}_i + n_e q_e \boldsymbol{v}_e + \varepsilon_o \frac{\partial \boldsymbol{E}}{\partial t})$$
$$\nabla \cdot \boldsymbol{B} = 0$$

$$p_{j} = C_{j} n_{j}^{\gamma} j = i, e$$
$$\frac{\partial n_{j}}{\partial t} + \nabla \cdot (n_{j} \boldsymbol{v}_{j}) = 0$$

 $\boldsymbol{j} = n_i q_i \boldsymbol{v}_i + n_e q_e \boldsymbol{v}_e$ $\rho_q = n_i q_i + n_e q_e$

1. Linearisation of plasma equations:

 $S = S_o + S_1$ S_o is the plasma variable in equilibrium

- S_1 is the perturbation of the plasma variable in equilibrium
- 2. the perturbation is a planar harmonic wave around the point of equilibrium S_0 :

 $S_1 = S_{10} e^{i(\mathbf{k} \cdot \mathbf{r} - \omega t)}$

- 3. Solve the linearised equations of the plasma to find the <u>dispersion relation</u> $\omega(\kappa)$:
- 4. We ignore the influence of the oscillations on the unperturbed quantities i.e. S_o):

$$m_j n_j \left(\frac{\partial \boldsymbol{v}_j}{\partial t} + (\boldsymbol{v}_j \cdot \boldsymbol{\nabla}) \boldsymbol{v}_j\right) = q_j n_j \left(\boldsymbol{E} + \boldsymbol{v}_j \times \boldsymbol{B}\right) - \boldsymbol{\nabla} \boldsymbol{p}_j \quad j=i,e$$

Plasma - One Fluid Equations - MHD

$$m_{j}n_{j}\left(\frac{\partial \boldsymbol{v}_{j}}{\partial t}+(\boldsymbol{v}_{j}\cdot\boldsymbol{\nabla})\boldsymbol{v}_{j}\right)=q_{j}n_{j}\left(\boldsymbol{E}+\boldsymbol{v}_{j}\times\boldsymbol{B}\right)-\boldsymbol{\nabla}\boldsymbol{p}_{j}$$

(

of fluid:

Current density:
$$J(\mathbf{r},t) = e(n_i v_i - n_e v_e)$$

Charge density: $\rho_q(\mathbf{r},t) = e(n_i - n_e)$
Mass density: $\rho_m(\mathbf{r},t) = (n_e m_e + n_i m_i) \cong n_i m_i$
because $m_e << m_i$ $n_e \cong n_i$
Pressure: $p = p_e + p_i$
Mass centre vel.
of fluid: $v = (m_i n_i v_i + m_e n_e v_e) / \rho_m$
 $J = n_i q_i v_i + n_e q_e v_e$
 $\rho_q = n_i q_i + n_e q_e$
For electrically neutral plasma
 $q_e n_e \sim q_i n_i$
For isotropic plasma:
 $p_j = n_j m_j v_j^2 = n_j K_B T_j$

$$\boldsymbol{J} = n_i q_i \boldsymbol{v}_i + n_e q_e \boldsymbol{v}_e$$
$$\rho_q = n_i q_i + n_e q_e$$

 $p_i = C_i n_i^{\gamma} j = i, e$

 $\frac{\partial n_{j}}{\partial t} + \boldsymbol{\nabla} \cdot (n_{j}\boldsymbol{v}_{j}) = 0$

ally neutral plasma:


MHD Equations













Ideal Magneto-Hydro Dynamics (Ideal MHD)

If we also consider that there is no accumulation of space charge i.e. $\rho_q=0$, the Simplified MHD equations including Maxwell equations are:



Closed system of equations – Good luck!!





Sound waves in a fluid: ion acoustic waves in a plasma

$$\rho_m(\frac{\partial \boldsymbol{v}}{\partial t} + (\boldsymbol{v} \cdot \boldsymbol{\nabla})\boldsymbol{v}) = -\boldsymbol{\nabla}\boldsymbol{p} = -\gamma \frac{p}{\rho_m} \boldsymbol{\nabla}\rho_m$$

Navier-Stokes – no viscosity

$$\frac{\nabla p}{p} = \gamma \frac{\nabla \rho_m}{\rho_m}$$

 $\frac{\partial \rho_m}{\partial t} + \nabla \cdot (\rho_m \upsilon) = 0$ Continuity equ

uation
ideal gas:
$$p = \frac{\rho_m KT}{m}$$

Linearising about a stationary equilibrium with uniform p_o and ρ_o : (keep only 1st order

$$\rho_{m} = \rho_{mo} + \rho_{1mo}e^{i(k\cdot r - \omega t)}$$

$$v = v_{o} + v_{10}e^{i(k\cdot r - \omega t)}, v_{o} = 0$$

$$Fluid is immobile and we investigate the vibrations of its velocity$$

$$-i\omega\rho_{1} + \rho_{o}i\mathbf{k} \cdot \mathbf{v}_{1} = 0$$

$$Continuity equation$$

$$\mathbf{k} = k\mathbf{x}, and \mathbf{v} = v\mathbf{x}$$

$$-i\omega\rho_{o}\mathbf{v}_{1} = -\gamma \frac{p_{o}}{\rho_{o}}i\mathbf{k}\rho_{1}$$
Navier-Stokes
$$Eliminating \rho_{1}$$

Eliminating ρ_1 :

$$-i\omega\rho_{mo}\boldsymbol{v}_{1} = -\gamma \frac{p_{o}}{\rho_{mo}}i\boldsymbol{k}\frac{\rho_{mo}i\boldsymbol{k}\boldsymbol{v}_{1}}{i\omega} \Leftrightarrow \omega^{2}\boldsymbol{v}_{1} = k^{2}\gamma \frac{p_{o}}{\rho_{mo}}\boldsymbol{v}_{1} \Leftrightarrow$$

$$p = \frac{\rho_{mo}KT}{m}$$

 $\frac{\omega}{k} = \left(\gamma \frac{p_o}{\rho_{mo}}\right)^{1/2} = \left(\gamma \frac{KT}{m}\right)^{1/2} \equiv c_s$ Sound waves velocity in a neutral gas. These are pressure waves which propagate through collisions among the air molecules.

In a plasma: Ion acoustic waves

Low frequency oscillations because of the large mass of the ion.

Ion fluid equation:

$$m_i n(\frac{\partial \boldsymbol{v}_i}{\partial t} + (\boldsymbol{v}_i \cdot \boldsymbol{\nabla})\boldsymbol{v}_i) = en\boldsymbol{E} - \boldsymbol{\nabla} p = -en\boldsymbol{\nabla} \phi - \gamma_i K T_i \boldsymbol{\nabla} n$$

For the electrons we assume m_e=0 Low frequency so we use plasma approximation, so:

$$n_i = n_e = n$$

Linearising (assuming plane waves): $-i\omega m_i n_o v_{i1} = -eik\phi_1 - \gamma_i KT_i ikn_1$

Ion sound waves in a plasma

The balance of forces on electrons requires:

$$n_e = n = n_o e^{\frac{e\phi_1}{KT_e}} = n_o \left(1 + \frac{e\phi_1}{KT_e} + \dots\right)$$

Since $E_o = 0$, $\varphi_o = 0$:

Boltzmann relation for electrons

$$n_e = n_{oe} e^{\frac{e\phi}{KT_e}}$$

Electrons are very mobile that their heat conductivity is huge! So we can consider isothermal electrons i.e. $\gamma=1$

The perturbation in the electron and therefore the ion density can be written:

$$n_1 = n_o \frac{e\phi_1}{KT_e}$$

Linearising the ion equation of continuity:

$$\begin{split} \frac{\partial n_1}{\partial t} + \boldsymbol{\nabla} \cdot (n_o \boldsymbol{v}_{i1} + n_1 \boldsymbol{v}_{i1}) &= 0 \Leftrightarrow \frac{\partial n_1}{\partial t} + n_o \boldsymbol{\nabla} \cdot \boldsymbol{v}_{i1} + \boldsymbol{v}_{i1} \cdot \boldsymbol{\nabla} n_o) = 0 \Leftrightarrow \\ \Leftrightarrow -i\omega n_1 &= -n_o i k \boldsymbol{v}_{i1} \end{split}$$

Ion sound waves in a plasma

So the set of equations are:

$$-i\omega m_{i}n_{o}\upsilon_{i1} = -eik\phi_{1} - \gamma_{\iota}KT_{i}ikn_{1}$$
$$n_{1} = n_{o}\frac{e\phi_{1}}{KT_{e}}$$

Substituting for ϕ_1 and n_1 in the equation of motion:

$$i\omega m_{i}n_{o}\upsilon_{i1} = en_{o}ik\frac{KT_{e}}{en_{o}} + \gamma_{\iota}KT_{i}ik\frac{n_{o}ik\upsilon_{i1}}{i\omega} \Leftrightarrow \omega^{2} = k^{2}\left(\frac{KT_{e} + \gamma_{\iota}KT_{\iota}}{m_{i}}\right)$$

$$\frac{\omega}{k} = \left(\frac{KT_e + \gamma_\iota KT_\iota}{m_i}\right)^{1/2} \equiv \upsilon_s$$

Ions suffer 1-D compression due to the plane wave we can use $\gamma_1=3$. Electrons are so fast that they have time to equalise their temperature everywhere, so they are isothermal and $\gamma_e=1$.

Dispersion relation of ion acoustic wave $-v_s$ is the sound speed in a plasma

Electron waves in a plasma

B=0 i.e. non magnetised plasma $T_e=T_i=0$ i.e. no thermal motion, ions are fixed in space and have uniform distribution, plasma is infinite in space.

Let's display electrons from the uniform ion background: E-field will be generated to restore neutrality.

Equation of motion for electrons:

$$m_e n_e (\frac{\partial \boldsymbol{v}_e}{\partial t} + (\boldsymbol{v}_e \cdot \boldsymbol{\nabla})\boldsymbol{v}_e) = -en_e \boldsymbol{E}$$

$$\varepsilon_o \nabla \cdot \boldsymbol{E} = \varepsilon_o \frac{\partial E}{\partial x} = e(n_i - n_e)$$

Continuity equation:

$$\frac{\partial n_e}{\partial t} + \boldsymbol{\nabla} \cdot (n_e \boldsymbol{v}_e) = 0$$

We chose 1D for simplicity

Linearisation:

$$n_e = n_o + n_1 \qquad \mathcal{U}_e = \mathcal{U}_o + \mathcal{U}_1 \qquad E = E_o + E_1$$

 $v_1 = v_{10}e^{i(kx-\omega t)}$ $n_1 = n_{10}e^{i(kx-\omega t)}$ $E = E_0e^{i(kx-\omega t)}$

Electron waves in a plasma

Initial conditions (uniform neutral plasma): In equilibrium: $n_{io}=n_{eo}$ and $n_{i1}=0$

$$\nabla n_{o} = v_{o} = E_{o} = 0 \qquad \frac{\partial}{\partial t} n_{o} = \frac{\partial}{\partial t} v_{o} = \frac{\partial}{\partial t} E_{o} = 0$$

$$m_{e} n_{e} (\frac{\partial v_{e}}{\partial t} + (v_{e} \cdot \nabla) v_{e}) = -en_{e} E \implies m_{e} (\frac{\partial v_{1}}{\partial t} + (v_{1} \cdot \nabla) v_{1}) = -eE$$

$$0 \text{ because quadratic}$$

$$\frac{\partial n_{e}}{\partial t} + \nabla \cdot (n_{e} v_{e}) = 0 \implies \frac{\partial n_{1}}{\partial t} + \nabla \cdot (n_{o} v_{1} + n_{1} v_{1}) = 0$$

$$0 \text{ because quadratic}$$

$$From \text{ Poisson's equation:}$$

$$\varepsilon_{o} \frac{\partial E_{1}}{\partial x} = -en_{1} \text{ Since } n_{0} = n_{eo} \text{ and } n_{1} = 0 \text{ by the assumption of fixed ions}$$

Electron waves in a plasma

So there are three equations and three unknowns, linearising:

$$\begin{aligned} -i\omega m_e \upsilon_1 &= -eE_1 \\ -i\omega n_1 &= -n_o ik\upsilon_1 \\ ik\varepsilon_o E_1 &= -en_1 \end{aligned}$$
Eliminating n_1 and E_1

$$-i\omega m_e \upsilon_1 &= -e\frac{-e}{ik\varepsilon_o} \frac{-n_o ik\upsilon_1}{-i\omega} &= -i\frac{n_o e^2}{\varepsilon_o \omega}\upsilon_1 \\ &\Rightarrow \omega^2 &= \frac{n_o e^2}{m_e \varepsilon_o} \text{ rad/sec} \end{aligned}$$
Characteristic plasma
frequency

$$\begin{aligned} \omega_{pe} &= f_{pe} \sim 9\sqrt{n_e} \\ \omega_{pe} &= 10^{18} m^{-3} \end{aligned}$$

$$\begin{aligned} \omega_{pe} &= 28GHz / Tesla \\ f_{pe} &\sim f_{ce} \text{ for } B \sim 0.32Tesla \text{ and } n \sim 10^{18} m^{-3} \end{aligned}$$



Thermal electron waves in a plasma



$$\nabla p_e = 3KT_e \nabla n_e$$

$$\omega^2 = \omega_p^2 + \frac{3}{2}k^2\upsilon_{th}^2$$







Magnetised plasmas

The existence of an external B_0 adds one more parameter to the cases we have studied i.e.:

Electrostatic waves ($B_1=0$) of low or high frequencies, Electromagnetic waves ($B_1\neq 0$) of low or high frequencies,

The new parameter is now whether the propagation is:

- 1) Perpendicular to \boldsymbol{B}_{0} : $\boldsymbol{k} \cdot \boldsymbol{B}_{0} = 0$
- 2) Parallel to \mathbf{B}_{o} : $\mathbf{k} \times \mathbf{B}_{o} = 0$

Note that waves travelling parallel to the perturbation E_1 are longitudinal i.e. $k \times E_1 = 0$ while waves travelling perpendicular to the perturbation E_1 are transverse i.e. $k \cdot E_1 = 0$

Home work: Show that all electrostatic waves are longitudinal and vise versa and that all transverse waves are electromagnetic but the opposite is not true.

High Frequency Electrostatic waves propagating transverse to external B field Upper Hybrid Frequency (UHF)

We will apply the same approaches as in non magnetised plasmas, "cold plasma", $T_e=T_i=0$ i.e. no thermal motion (thus $P_e=P_i=0$) ions are fixed in space and have uniform distribution, plasma is infinite in space. Since the waves have high frequency, ions can be considered immobile because of their large mass.

Since electrostatic waves are longitudinal, $\mathbf{k} \times \mathbf{E}_1 = 0$ and we also consider that $\mathbf{k} \cdot \mathbf{B}_0 = 0$

To identify the dispersion relation the following equations are needed:

$$m_{e}n_{e}\left(\frac{\partial \boldsymbol{v}_{e}}{\partial t} + (\boldsymbol{v}_{e}\cdot\boldsymbol{\nabla})\boldsymbol{v}_{e}\right) = -en_{e}\left[\boldsymbol{E}_{1} + \boldsymbol{v}_{e}\times\boldsymbol{B}_{o}\right] \qquad \frac{\partial n_{e}}{\partial t} + \boldsymbol{\nabla}\cdot(n_{e}\boldsymbol{v}_{e}) = 0$$
$$\varepsilon_{o}\boldsymbol{\nabla}\cdot\boldsymbol{E}_{1} = e(n_{o}-n_{e}) \qquad n_{io} = n_{eo} = n_{o}$$

High Frequency Electrostatic waves propagating transverse to external B field Upper Hybrid Frequency (UHF) E_x

Lets consider that $\mathbf{E}_1 = \mathbf{E}_1 \mathbf{x}$, $\mathbf{k} = k\mathbf{x}$, $\mathbf{B}_0 = \mathbf{B}_0 \mathbf{z}$ In order for the eq. of motion to be obeyed, v_e should lye on the (x,y) plane

Linearising the three equations (motion, continuity, Poisson) and keeping first order terms:

$$-i\omega m_e v_{e1x} = -eE_1 - ev_{e1y}B_0 \qquad -i\omega m_e v_{e1y} = ev_{e1x}B_0$$
$$-i\omega n_{e1} + ikn_o v_{e1x} = 0 \qquad ik\varepsilon_o E_1 = -en_{e1}$$

From the first two (eq. of motion):

Substituting in the third (continuity) and using the fourth (Poisson):

$$\upsilon_{e1x} = \frac{eE_1}{im_e \omega \left(1 - \frac{\Omega_e^2}{\omega_e^2}\right)} \quad \Omega_e = \frac{eB_o}{m_e}$$

$$\left(-\frac{\omega^2}{\omega_e^2} + 1 + \frac{\Omega_e^2}{\omega_e^2}\right) E_1 = 0 \Leftrightarrow \omega^2 = \omega_e^2 + \Omega_e^2 \equiv \omega_{UH}^2$$

High Frequency Electrostatic waves propagating transverse to external B field Upper Hybrid Frequency (UHF)

The electrons perform an additional cyclotron oscillation due to the presence of the external magnetic field.

Group velocity is zero as long as thermal motion is neglected.

Electrostatic electron wave along B_o are the usual plasma oscillations with $\omega = \omega_p$

Home work 1: Why in the above analysis we ignored the last three Maxwell's equations (Gauss's law for magnetism, Ampere's law, Faraday's law) and we used only the Poisson's equation.

Low Frequency Electrostatic waves propagating transverse to external B field Lower Hybrid Frequency (LHF)

Since electrostatic waves are investigated, only Poisson's equation could be used From the Maxwell's equations, but since frequency is low ions can follow the vibrations of electrons and $n_{e1}=n_{i1}$. So we have to use the equations of motion and continuity for both ions and electrons, thus the equations which describe the propagation of such waves are:

$$m_{e}n_{e}\left(\frac{\partial \boldsymbol{v}_{e}}{\partial t} + (\boldsymbol{v}_{e}\cdot\boldsymbol{\nabla})\boldsymbol{v}_{e}\right) = -en_{e}\left[\boldsymbol{E}_{1} + \boldsymbol{v}_{e}\times\boldsymbol{B}_{o}\right] \qquad \frac{\partial n_{e}}{\partial t} + \boldsymbol{\nabla}\cdot(n_{e}\boldsymbol{v}_{e}) = 0$$
$$m_{i}n_{i}\left(\frac{\partial \boldsymbol{v}_{i}}{\partial t} + (\boldsymbol{v}_{i}\cdot\boldsymbol{\nabla})\boldsymbol{v}_{i}\right) = -en_{i}\left[\boldsymbol{E}_{1} + \boldsymbol{v}_{i}\times\boldsymbol{B}_{o}\right] \qquad \frac{\partial n_{i}}{\partial t} + \boldsymbol{\nabla}\cdot(n_{i}\boldsymbol{v}_{i}) = 0$$
$$n_{io} = n_{eo} = n_{o}$$

After linearisation:

Low Frequency Electrostatic waves propagating transverse to external B field Lower Hybrid Frequency (LHF)

$$m_{e}n_{o}\frac{\partial \boldsymbol{v}_{e}}{\partial t} = -\gamma_{e}k_{B}T_{e}\boldsymbol{\nabla}n_{e1} - en_{o}\left[\boldsymbol{E}_{1} + \boldsymbol{v}_{e} \times \boldsymbol{B}_{o}\right] \qquad \frac{\partial n_{e1}}{\partial t} + n_{o}\boldsymbol{\nabla} \cdot \boldsymbol{v}_{e} = 0$$

$$m_{i}n_{o}\frac{\partial \boldsymbol{v}_{i}}{\partial t} = -\gamma_{i}k_{B}T_{i}\boldsymbol{\nabla}n_{e1} + en_{o}\left[\boldsymbol{E}_{1} + \boldsymbol{v}_{i} \times \boldsymbol{B}_{o}\right] \qquad \frac{\partial n_{e1}}{\partial t} + n_{o}\boldsymbol{\nabla} \cdot \boldsymbol{v}_{i} = 0$$

$$\nabla P_{j} = \gamma_{j}k_{B}T_{j}\boldsymbol{\nabla}n_{j}$$

where we have used that $n_{e1}=n_{i1}$ As previously $\mathbf{B}_{o}=\mathbf{B}_{o}\mathbf{z}$, and the waves propagate solutions ~ $e^{i(\mathbf{k}\cdot\mathbf{r}-\omega t)}$ on the (x,z) plane parallel with the perturbation of the E field, $\mathbf{k}//\mathbf{E}_{1}$

Add by parts:

$$-i\omega n_o(m_e \boldsymbol{v}_e + m_i \boldsymbol{v}_i) = -i\boldsymbol{k} n_{e1}(\gamma_e k_B T_e + \gamma_i k_B T_i) + en_o(\boldsymbol{v}_i - \boldsymbol{v}_e) \times \boldsymbol{B}_o$$

Low Frequency Electrostatic waves propagating transverse to external B field Lower Hybrid Frequency (LHF)

From the continuity equations : $\mathbf{k} \cdot \mathbf{v}_{e_1} = \mathbf{k} \cdot \mathbf{v}_{i_1} = \frac{\omega n_{e_1}}{n_o}$ So equation of motion becomes:

$$-i\omega^2 n_o(m_e + m_i)\frac{n_{e1}}{n_o} = -ik^2 n_{e1}(\gamma_e k_B T_e + \gamma_i k_B T_i) + en_o \mathbf{k} \cdot [(\mathbf{v}_i - \mathbf{v}_e) \times \mathbf{B}_o]$$
(1)

$$\boldsymbol{k} \cdot [(\boldsymbol{v}_i - \boldsymbol{v}_e) \times \boldsymbol{B}_o] = k_x [(\boldsymbol{v}_{iy} - \boldsymbol{v}_{ey}) \boldsymbol{B}_o]$$

But to find the dispersion relation from this equation we need to express the velocities as functions of the density.

For this reason we externally multiply the electron equation of motion (two slides before) with \mathbf{k} and we take:

$$-i\omega m_{e}n_{o}(\mathbf{k} \times \mathbf{v}_{e}) = -en_{o}\mathbf{k} \times (\mathbf{v}_{e} \times \mathbf{B}_{o}) - \begin{bmatrix} \upsilon_{ex} = -\frac{i\omega}{\Omega_{e}}\upsilon_{ey} \\ \upsilon_{ez} = +i\upsilon_{ey}\frac{k_{z}}{k_{x}}\frac{\Omega_{e}}{\omega}\left(1-\frac{\omega^{2}}{\Omega_{e}^{2}}\right) \end{bmatrix}$$

Low Frequency Electrostatic waves propagating transverse to external B field Lower Hybrid Frequency (LHF)

Electron continuity equations becomes:

$$k_{x}\upsilon_{ex} + k_{z}\upsilon_{ez} = \omega \frac{n_{e1}}{n_{o}} \Leftrightarrow \omega \frac{n_{e1}}{n_{o}} = \left[k_{x}\left(-\frac{i\omega}{\Omega_{e}}\right) + k_{z}\left(\frac{+i\Omega_{e}}{\omega}\right)\frac{k_{z}}{k_{x}}\left(1 - \frac{\omega^{2}}{\Omega_{e}^{2}}\right)\right]\upsilon_{ey}$$

By substituting v_{ey} in the previous relations for v_{ex} , v_{ez} we have written v_{ex} , v_{ez} as Functions of n_{e1} . Substituting in eq (1) of the previous slide and considering $m_e << m_{i}$, we obtain the dispersion relation:

$$1 - \frac{k^2 c_s^2}{\omega^2} + \frac{\Omega_i}{\omega} \left[\frac{1}{-\frac{\omega}{\Omega_e} + \frac{\Omega_e}{\omega} \frac{k_z^2}{k_x^2} \left(1 - \frac{\omega^2}{\Omega_e^2}\right)} - \frac{1}{\frac{\omega}{\Omega_i} - \frac{\Omega_i}{\omega} \frac{k_z^2}{k_x^2} \left(1 - \frac{\omega^2}{\Omega_i^2}\right)} \right] = 0$$
$$c_s^2 = \frac{\gamma_e k_B T_e + \gamma_i k_B T_i}{m_i}$$
General dispersion relation

Low Frequency Electrostatic waves propagating transverse to external B field Lower Hybrid Frequency (LHF)

If for instance $k_x=0$ (k//B_o) we obtain : $\omega^2 = c_s^2 k_z^2$ as we have previously calculated for B_o=0 (ιοντακουστικές ταλαντώσεις).

If k_z=0 the dispersion relation is
$$1 - \frac{k_x^2 c_s^2}{\omega^2} - \frac{\Omega_i \Omega_e}{\omega^2} - \frac{\Omega_i^2}{\omega^2} = 0$$

Because $\Omega_i^2 \ll |\Omega_i \Omega_e|$ we have: $\omega^2 = c_s^2 k_x^2 + \Omega_e \Omega_i$

When $k_x \rightarrow 0$ we have: $\omega_{LH}^2 = \Omega_i \Omega_e$ This is the Lower Hybrid Frequency – κάτω υβριδική συχνότητα

High Frequency Electromagnetic waves propagating transverse to external B field

Perpendicular propagation $\mathbf{k} \perp \mathbf{B}_o$, if we take transverse waves with $\mathbf{k} \perp \mathbf{E}_1$ there are two choices, 1) \mathbf{E}_1 can be parallel to \mathbf{B}_0 and 2) \mathbf{E}_1 can be perpendicular to \mathbf{B}_0

Ordinary wave $\mathbf{k} \perp \mathbf{B}_{o}, \mathbf{E} \parallel \mathbf{B}_{o}$

For \mathbf{E}_1 parallel to \mathbf{B}_0 we can take $\mathbf{B}_0 = \mathbf{B}_0 \mathbf{z}$, $\mathbf{E}_1 = \mathbf{E}_1 \mathbf{z}$, $\mathbf{k} = \mathbf{k} \mathbf{x}$ The wave equation is (as in the case for EM waves with $\mathbf{B}_0 = 0$)

$$\left(\omega^2 - c^2 k^2\right) \boldsymbol{E}_1 = \frac{-i\omega \boldsymbol{j}_1}{\varepsilon_o} = \frac{in_o e\omega \boldsymbol{v}_{e1}}{\varepsilon_o} \quad \boldsymbol{j}_1 = -en_{oe} \boldsymbol{v}_{e1} \quad n_{oe} = n_o$$

Because $E_1 = E_1 z$, we need only the v_{ez} component which is given by the particle equation of motion

 $m_e \frac{\partial v_{ez}}{\partial t} = -eE_z$ Since everything is the same as the equation with B=0 remember

$$\boldsymbol{v}_{e1} / \boldsymbol{E}_{1} \Rightarrow \boldsymbol{v}_{e1} \times \boldsymbol{B}_{o} = 0$$

$$\omega^2 = \omega_{pe}^2 + c^2 k^2$$

So the O-wave does not feel the existence of the external B_0

High Frequency Electromagnetic waves propagating transverse to external B field

Perpendicular propagation $\mathbf{k} \perp \mathbf{B}_o$, transverse waves $\mathbf{k} \perp \mathbf{E}_1$, \mathbf{E}_1 perpendicular to \mathbf{B}_o

Extra-ordinary wave $k \perp B_o, E \perp B_o$

When E_1 is perpendicular to B_0 , the electron motion will be affected by B_0 since $v \times B_0$ force is not zero as for the previous case (ordinary wave)

To treat this case we can take $\mathbf{B}_0 = \mathbf{B}_0 \mathbf{z}$, $\mathbf{k} = \mathbf{k}\mathbf{x}$. However, for the \mathbf{E}_1 someone should allow for the development of one more component along \mathbf{k} , so even if we start with $\mathbf{E}_1 = \mathbf{E}_1 \mathbf{y}$ the \mathbf{E}_1 for generality can be $\mathbf{E}_1 = \mathbf{E}_x \mathbf{x} + \mathbf{E}_y \mathbf{y}$ (Actually, it turns out that the wave will become elliptically polarised instead of plane (linearly) polarised, so it will become **partly longidudinal** and **partly transverse**.



High Frequency Electromagnetic waves propagating transverse to external B field

The system of equations is:

$$m_{e}n_{e}\left(\frac{\partial \boldsymbol{v}_{e}}{\partial t} + (\boldsymbol{v}_{e}\cdot\boldsymbol{\nabla})\boldsymbol{v}_{e}\right) = -en_{e}\left[\boldsymbol{E} + \boldsymbol{v}_{e}\times\boldsymbol{B}\right] \qquad \frac{\partial n_{e}}{\partial t} + \boldsymbol{\nabla}\cdot\left(n_{e}\boldsymbol{v}_{e}\right) = 0$$
$$\boldsymbol{\nabla}\times\boldsymbol{E} = -\frac{\partial \boldsymbol{B}}{\partial t} \qquad \boldsymbol{\nabla}\times\boldsymbol{B} = \mu_{o}\left(\boldsymbol{j} + \varepsilon_{o}\frac{\partial \boldsymbol{E}}{\partial t}\right) \qquad \boldsymbol{j} = -en_{e}\boldsymbol{v}_{e}$$

The linearised system of equations is (linearised as in previous cases):

$$m_{e}n_{o}\frac{\partial \boldsymbol{v}_{e1}}{\partial t} = -en_{o}\boldsymbol{E}_{1} - n_{o}\boldsymbol{v}_{e1} \times \boldsymbol{B}_{o} \qquad \boldsymbol{\nabla} \times \boldsymbol{E}_{1} = -\frac{\partial \boldsymbol{B}_{1}}{\partial t}$$
$$\boldsymbol{\nabla} \times \boldsymbol{B}_{1} = \mu_{o}(\boldsymbol{j} + \varepsilon_{o}\frac{\partial \boldsymbol{E}_{1}}{\partial t}) \qquad \boldsymbol{j} = -en_{o}\boldsymbol{v}_{e1}$$

$$n_{e} = n_{eo} + n_{e1} \qquad v_{e} = v_{eo} + v_{e1} \qquad E = E_{1} \qquad B = B_{o} + B_{1} \qquad n_{eo} = n_{o}$$

$$n_{e1} = n_{e10}e^{i(kx-\omega t)} \qquad v_{e1} = v_{e10}e^{i(kx-\omega t)} \qquad E_{1} = E_{01}e^{i(kx-\omega t)} \qquad B_{1} = B_{01}e^{i(kx-\omega t)} \qquad v_{eo} = 0$$

High Frequency Electromagnetic waves propagating transverse to external B field The three linearised equations become:

$$-i\omega m_e v_x = -eE_x - ev_y B_o$$

$$-i\omega m_e v_y = -eE_y - ev_x B_o$$

$$ikE_y = i\omega B_1$$

$$-ikB_1 = -en_o v_y - i\omega E_y$$

$$0 = -en_o v_x - i\omega E_x$$

From the last three:

$$\upsilon_{x} = \frac{e}{m\omega} \left(-iE_{x} - \frac{\omega_{c}}{\omega} E_{y} \right) \left(\frac{1}{1 - \frac{\omega_{c}^{2}}{\omega^{2}}} \right)$$
$$\upsilon_{y} = \frac{e}{m\omega} \left(-iE_{y} - \frac{\omega_{c}}{\omega} E_{x} \right) \left(\frac{1}{1 - \frac{\omega_{c}^{2}}{\omega^{2}}} \right)$$

Replacing into the first two and taking into account that:

$$\omega_{pe}^2 = \frac{n_e e^2}{\varepsilon_o m_e}$$

High Frequency Electromagnetic waves propagating transverse to external B field

$$\left[\begin{array}{ccc} \omega^{2} \left(1 - \frac{\omega_{c}^{2}}{\omega^{2}}\right) - \omega_{pe}^{2} & i \frac{\omega_{pe}^{2} \omega_{c}}{\omega} \\ \left(\omega^{2} - c^{2} k^{2}\right) \left(1 - \frac{\omega_{c}^{2}}{\omega^{2}}\right) - \omega_{pe}^{2} & -i \frac{\omega_{pe}^{2} \omega_{c}}{\omega} \end{array} \right] \left[\begin{array}{c} E_{x} \\ E_{y} \end{array} \right] = \left[\begin{array}{c} 0 \\ 0 \end{array} \right]$$

For E_x , E_y other than zero this system has a solution only if the determinant is zero

Taking into account that: $\omega_{pe}^2 + \omega_c^2 = \omega_{UH}^2 = \omega_h^2$

$$\left(\omega^{2}-\omega_{h}^{2}\right)\left[\omega^{2}-\omega_{h}^{2}-c^{2}k^{2}\left(1-\frac{\omega_{c}^{2}}{\omega^{2}}\right)\right] = \left(\frac{\omega_{pe}^{2}\omega_{c}}{\omega}\right)^{2}$$
$$\frac{\omega^{2}-\omega_{h}^{2}-\left[\frac{\left(\omega_{pe}^{2}\omega_{c}\right)^{2}}{\omega^{2}\left(\omega^{2}-\omega_{h}^{2}\right)}\right]}{\omega^{2}\left(\omega^{2}-\omega_{h}^{2}\right)}\right]}{\omega^{2}-\omega_{c}^{2}}$$
Repared with

Replacing back the $(\omega_h)^2$ and multiplying through with $\omega^2 - \omega_h^2$ we find:

High Frequency Electromagnetic waves propagating transverse to external B field

$$\frac{c^2k^2}{\omega^2} = \frac{c^2}{\upsilon_{\varphi}^2} = 1 - \frac{\omega_{pe}^2}{\omega^2} \frac{\omega^2 - \omega_{pe}^2}{\omega^2 - \omega_h^2}$$

Dispersion relation for the extraordinary wave (X-wave

Cutoff and Resonance frequencies for the X-wave

For $k \rightarrow \infty$ $\omega \rightarrow \omega_h$ so that a **resonance** occurs at a point in the plasma where:

$$\omega_h^2 = \omega_{pe}^2 + \omega_c^2$$

This is the dispersion relation for electrostatic waves propagating across \mathbf{B}_{o} . As the wave approaches the resonant point both its phase velocity and its group velocity approach zero and the energy is converted into upper hybrid oscillations. The X-wave is partly electromagnetic and partly electrostatic, so at resonance this wave losses its electromagnetic character and becomes an electrostatic oscillation.

High Frequency Electromagnetic waves propagating transverse to external B field

The cutoffs of the X-wave are found by setting k equal to zero at the dispersion relation. After some easy algebra we conclude to the following equation:

$$\left(1 - \frac{\omega_{pe}^2}{\omega^2}\right)^2 = \frac{\omega_c^2}{\omega^2} \Leftrightarrow 1 - \frac{\omega_{pe}^2}{\omega^2} = \pm \frac{\omega_c}{\omega} \Leftrightarrow \omega^2 \mp \omega \omega_c - \omega_{pe}^2 = 0 \Leftrightarrow$$

$$\Leftrightarrow \left\{ \begin{array}{l} \omega_{R} = \frac{1}{2}\omega_{c} + \left(\omega_{pe}^{2} + \frac{1}{4}\omega_{c}^{2}\right)^{1/2} \\ \omega_{L} = -\frac{1}{2}\omega_{c} + \left(\omega_{pe}^{2} + \frac{1}{4}\omega_{c}^{2}\right)^{1/2} \end{array} \right.$$

High Frequency Electromagnetic waves propagating transverse to external B field



- ✓ From the dispersion relation it is clear that at $\omega = \omega_{pe}$ the wave propagates at the velocity of light.
- ✓ For $\omega < \omega_L$ there is another region of non propagation.

The cutoff and resonance frequencies divide the dispersion diagram into regions of propagation and non propagation.

- At large ω (or low density) the phase velocity approaches the velocity of light.
- As the wave travels further the phase velocity increases until the right hand cutoff $\omega = \omega_R$ is met. There the phase velocity becomes infinite.
- ✓ Between $\omega = \omega_R$ and $\omega = \omega_h$ the $(\upsilon_{\phi})^2$ is negative and propagation is prohibited.
- ✓ At $\omega = \omega_h$ there is a resonance and υ_{φ} goes to zero.
- ✓ Between $\omega = \omega_h$ and $\omega = \omega_L$ propagation is again allowed. In this region the wave travels either faster or slower than c depending on whether ω is smaller or larger than ω_{pe} .

High Frequency Electromagnetic waves propagating transverse to external B field

For the O-wave:



High Frequency Electromagnetic waves propagating parallel to external B field

Let's consider that k lies along z axis and allow E_1 to have both transverse components (general case)

Wave equation:
$$\nabla \times (\nabla \times E_{1}) = \nabla (\nabla \cdot E_{1}) - \nabla^{2} E_{1} = -\nabla \times \frac{\partial}{\partial t} B_{1}$$

 $(\omega^{2} - c^{2}k^{2})E_{x} = \frac{\omega_{p}^{2}}{1 - \frac{\omega_{c}^{2}}{\omega^{2}}} \left(E_{x} - \frac{i\omega_{c}}{\omega}E_{y}\right)$
 $\delta = \frac{\omega_{p}^{2}}{1 - \frac{\omega_{c}^{2}}{\omega^{2}}} \left(\omega^{2} - c^{2}k^{2} - \delta\right)E_{x} + i\delta\frac{\omega_{c}}{\omega}E_{y} = 0$
 $(\omega^{2} - c^{2}k^{2})E_{y} = \frac{\omega_{p}^{2}}{1 - \frac{\omega_{c}^{2}}{\omega^{2}}} \left(E_{y} + \frac{i\omega_{c}}{\omega}E_{x}\right)$
 $\left(\omega^{2} - c^{2}k^{2} - \delta\right)E_{y} - i\delta\frac{\omega_{c}}{\omega}E_{x} = 0$
Setting determinant equal to zero:

Setting determinant equal to zero:

remember

$$\left(\omega^2 - c^2 k^2 - \delta\right)^2 = \left(\frac{\delta\omega_c}{\omega}\right)^2 = 0 \iff \omega^2 - c^2 k^2 - \delta = \pm \frac{\delta\omega_c}{\omega}$$

High Frequency Electromagnetic waves propagating parallel to external B field

...k lies along z axis and allow E_1 to have both transverse components (general case)

$$\omega^{2} - c^{2}k^{2} - \delta = \pm \frac{\delta\omega_{c}}{\omega} \iff \omega^{2} - c^{2}k^{2} = \delta\left(1 \pm \frac{\omega_{c}}{\omega}\right) = \frac{\omega_{p}^{2}}{1 - \frac{\omega_{c}^{2}}{\omega^{2}}} \left(1 \pm \frac{\omega_{c}}{\omega}\right) = \omega_{p}^{2} \frac{1 \pm \frac{\omega_{c}}{\omega}}{\left[1 + \frac{\omega_{c}}{\omega}\right]\left[1 - \frac{\omega_{c}}{\omega}\right]} = \frac{\omega_{p}^{2}}{1 \mp \frac{\omega_{c}}{\omega}}$$

The \mp sign shows that two solutions of two different waves that can propagate along B_o exist. Their dispersion relations are:



High Frequency Electromagnetic waves propagating parallel to external B field

- (-) R-wave Right hand circular polarisation
- (+)L-wave Left hand circular polarisation

Since these equations depend only on k^2 the direction of rotation of the *E* vector is independent of the sign of *k* thus the polarisation is the same for waves propagating in the opposite direction.

Cutoff and Resonance frequencies for R and L waves

For the <u>**R** wave</u> $k \rightarrow \infty$ at $\omega = \omega_c$ the wave is therefore in **resonance** with the cyclotron motion of the electrons. The direction of rotation of the polarisation plane is the same as the direction of the gyration of electrons so the wave loses its energy in continuously accelerating the electrons and therefore cannot propagate.

The <u>L wave</u> does **not** have a cyclotron resonance with the electrons because it rotates in the opposite direction. Actually as seen from its dispersion relation the L wave does not have a resonance for positive ω (in some considerations ω can take negative values). (If we had included ions motion, the L wave would have a resonance at $\omega = \omega_{ci}$)

High Frequency Electromagnetic waves propagating parallel to external B field

Remember that in our convection ω is always positive and waves propagating in the – direction (i.e. –x) are described by negative *k*.

Cutoffs: For $k \rightarrow 0$

$$\omega_{R} = \frac{1}{2}\omega_{c} + \left(\omega_{pe}^{2} + \frac{1}{4}\omega_{c}^{2}\right)^{1/2}$$
$$\omega_{L} = -\frac{1}{2}\omega_{c} + \left(\omega_{pe}^{2} + \frac{1}{4}\omega_{c}^{2}\right)^{1/2}$$

- ✓ Same equations as for the cutoffs of the X-wave
- ✓ The R-wave (-) has the higher cutoff frequency ω_R while the L-wave (+) has the lower cutoff frequency

High Frequency Electromagnetic waves propagating parallel to external B field



The dispersion diagram for the R and L waves. Regions of non propagation are for $v_{\phi}^2/c^2 < 0$

- The L wave (+) has a stop band at low frequencies and it behaves like the O-wave except that the cutoff occurs at ω_L instead of ω_p
 - The R-wave (-) has a stop band between ω_R and ω_c , but there is a second band of propagation with $\upsilon_{\phi} < c$ bellow ω_c . The wave in this low frequency region is called the *"whistler mode"* ($\sigma \varphi \upsilon \rho \upsilon \chi \tau \dot{\alpha}$) and is of extreme importance for ionospheric phenomena

High Frequency Electromagnetic waves propagating parallel to external B field

Summary

The principal electromagnetic waves propagating along B_o are a Right-hand (R) and a Left-hand (L) circularly polarised wave, the principal waves propagating across B_o are a plane-polarised wave (O-wave) and an elliptically polarised wave (X-wave)

High Frequency Electromagnetic waves propagating parallel to external B field

Faraday rotation

From the previous diagram it is clear that for large ω , the R wave travels faster than the L wave. Consider the plane polarised wave to be the sum of an R wave and an L wave (of course at the same frequency).

After let's say N cycles, the E_L and E_R vectors will return to their initial positions. However, after propagating a given distance d in a plasma the R and L waves will have undergone a different number of cycles since they require a different amount of time to cover the distance.



A plane polarised wave is the sum of left and right Handed circularly polarised waves

Since the L wave travels more slowly (in a plasma) it will have undergone N+a cycles at the position where R has undergone N cycles

The plane of polarisation is rotated

High Frequency Electromagnetic waves propagating parallel to external B field

Faraday rotation

$$\theta(rad) = \frac{e^{3}\lambda_{o}^{2}}{8\pi^{2}m_{e}^{2}\varepsilon_{o}c^{3}}\int_{\ell}\frac{n_{e}B_{o}}{\sqrt{1-\frac{n_{e}}{n_{c}}}}dz \quad \omega >> \omega_{c}$$

4.14
4.23
4.24
Chen
$$\frac{e^{3}}{8\pi^{2}m_{e}^{2}\varepsilon_{o}c^{3}} = 2.6312 \times 10^{-13} (T^{-1})$$

$$\frac{e^{3}\lambda_{e}^{2}}{8\pi^{2}m_{e}^{2}\varepsilon_{o}c^{3}}\int_{\ell}n_{e}B_{o}dz \quad n_{c} >> n_{e}$$

$$n_{c} = \frac{\omega^{2} m_{e} \varepsilon_{o}}{e^{2}} \qquad \omega_{pe}^{2} = \frac{n_{e} e^{2}}{\varepsilon_{o} m_{e}} \qquad 1 - \frac{\omega_{pe}^{2}}{\omega^{2}} = 1 - \frac{n_{e}}{n_{c}}$$


Waves in plasmas



Electron waves (electrostatic):

$$B_{o} = 0$$

or
$$\omega^{2} = \omega_{p}^{2} + \frac{3}{2}k^{2}\upsilon_{thermal}^{2}$$

$$k / / B_{o}$$

 $\boldsymbol{k} \perp \boldsymbol{B}_{o} \qquad \qquad \boldsymbol{\omega}^{2} = \boldsymbol{\omega}_{p}^{2} + \boldsymbol{\omega}_{c}^{2}$

Plasma oscillations

Upper hybrid oscillations

Ion waves (electrostatic):

$$\boldsymbol{B}_{o} = 0$$

or
$$\boldsymbol{\omega}^{2} = k^{2} \boldsymbol{\upsilon}_{s}^{2} = k^{2} \left(\frac{\boldsymbol{\gamma}_{e} k_{B} T_{e} + \boldsymbol{\gamma}_{\iota} k_{B} T_{\iota}}{m_{i}} \right)^{1/2}$$

$$\boldsymbol{k} / \boldsymbol{\beta}_{o}$$

Acoustic waves

$$\boldsymbol{k} \perp \boldsymbol{B}_{o} < \begin{cases} \boldsymbol{\omega}^{2} = \boldsymbol{\Omega}_{c}^{2} + \boldsymbol{k}^{2} \boldsymbol{\upsilon}_{s}^{2} \\ \text{or} \\ \boldsymbol{\omega}^{2} = \boldsymbol{\omega}_{l}^{2} = \boldsymbol{\Omega}_{c} \boldsymbol{\omega}_{c} \end{cases}$$



Electrostatic ion cyclotron waves

Lower hybrid oscillations







Electron way		
$B_o = 0$	$\omega^2 = \omega_p^2 + k^2 c^2$	Light waves
$\boldsymbol{k} \perp \boldsymbol{B}_o, \boldsymbol{E}_1 / \boldsymbol{B}_o$	$\frac{c^2k^2}{\omega^2} = 1 - \frac{\omega_p^2}{\omega^2}$	O - wave
$\boldsymbol{k} \perp \boldsymbol{B}_o, \boldsymbol{E}_1 \perp \boldsymbol{B}_o$	$\frac{c^2k^2}{\omega^2} = 1 - \frac{\omega_p^2}{\omega^2} \frac{\omega^2 - \omega_p^2}{\omega^2 - \omega_h^2}$	X - wave
k / / B _o	$\int \frac{c^2 k^2}{\omega^2} = 1 - \frac{\omega_p^2}{\omega^2 - \omega \omega_c}$	R - wave whistler mode
	$\frac{c^2k^2}{\omega^2} = 1 - \frac{\omega_p^2}{\omega^2 + \omega\omega_c}$	L - wave









Ion waves (electromagnetic):		
$B_o = 0$	по	There is no electromagnetic wave
k / / B _o	$\omega^2 = k^2 v_A^2$	Alfvèn wave
$\boldsymbol{k} \perp \boldsymbol{B}_{o}$	$\frac{\omega^{2}}{k^{2}} = c^{2} \frac{v_{s}^{2} + v_{A}^{2}}{c^{2} + v_{A}^{2}}$	Magnetosonic wave

The above dispersion relations cover the main propagation geometries





